

Reminder on Multiplication Factor

- For a reactor with neutron production rate P , absorption rate (A) and leakages (L), we can define the multiplication factor as:

$$k = \text{Multiplication Factor} = \frac{\text{The number of fissions in generation } i}{\text{The number of fissions in previous generation } (i - 1)} = \frac{P}{A + L}$$

- If one fission leads to more than one fission the fission rate increases in time exponentially
 - Such a reactor is called **supercritical and $k > 1$**
- If one fission leads to less than one fission the fission rate decreases in time and eventually the chain reaction stops
 - Such a reactor is called **subcritical and $k < 1$**
- If one fission leads on average to exactly one other fission the chain reaction is self-sustained.
 - Such reactor is **called critical and $k = 1$**

- For a critical infinite system:

$$k_{\infty} = \frac{P}{A} = \frac{\bar{v}\Sigma_f}{\Sigma_a} = \frac{\bar{v}\Sigma_f}{\Sigma_{a,fuel} + \Sigma_{a,nonFuel}} = \frac{\bar{v}\Sigma_f}{\Sigma_{a,fuel}} \cdot \frac{\Sigma_{a,fuel}}{\Sigma_{a,fuel} + \Sigma_{a,nonFuel}} = \eta \cdot f = 1$$

where **f is the fuel utilization factor**, i.e. how many neutrons are absorbed in the fuel per total absorption in the system.

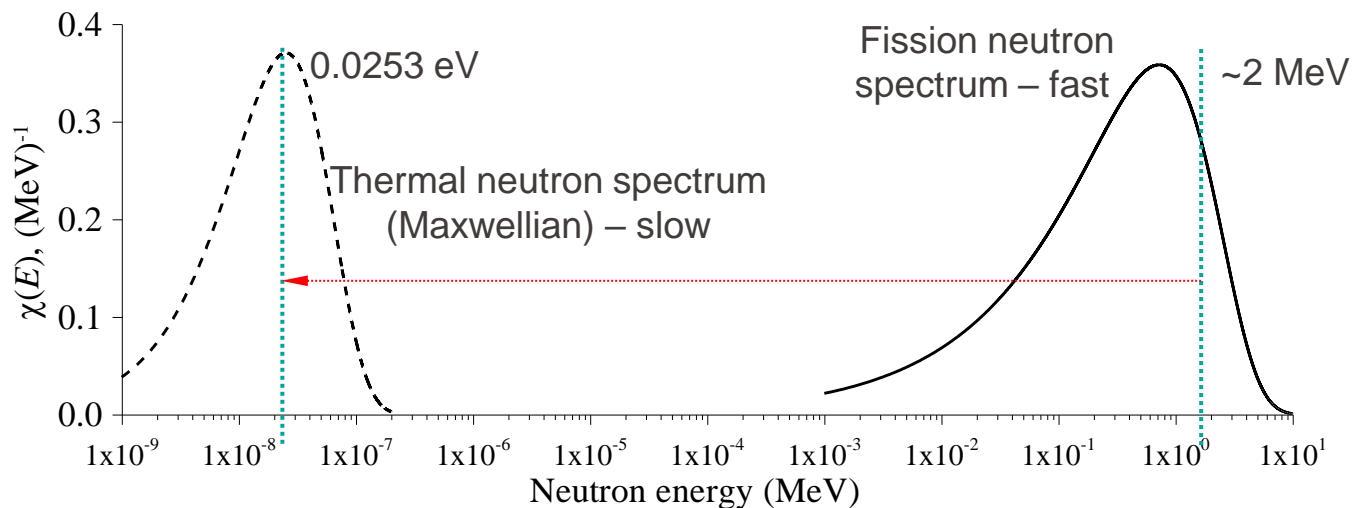
- In finite systems some neutrons escape, so $k_{\infty} = 1$ is not enough. One talks about k-effective

$$k_{\text{eff}} = \frac{P}{A + L} = \frac{P}{A} \cdot \frac{A}{A + L} = k_{\infty} \cdot P_{NL} = 1$$

where **P_{NL} is the non-leakage probability**, i.e. the probability that one neutron remains in the system without leaking and is finally absorbed in the system (parasitic or fissile).

Fission spectrum → Thermal spectrum

- For a thermal reactor, neutrons are born fast and are absorbed in thermal range.
- “Slowing down factor” of almost 10^8 ! (From approx. 2 MeV to 0.0253 eV)
- Moderators needed (light nuclei: H₂O, graphite,...)



For a thermal reactor one can express the multiplication factor in a convenient form:

$$k_{\infty} = \frac{P}{A} = \frac{P}{P_{th}} \times \frac{P_{th}}{A_{th}^{fuel}} \times \frac{A_{th}^{fuel}}{A_{th}} \times \frac{A_{th}}{A} = \varepsilon \times \eta_{th} \times f \times p$$

$$\varepsilon = \frac{P}{P_{th}} = \frac{P_{th} + P_{fast}}{P_{th}} = 1 + \frac{P_{fast}}{P_{th}}$$

Fast fission factor: bonus from fast fissions in non-fissile material. Typical value ~ 1.04

$$\eta_{th} = \frac{P_{th}}{A_{th}^{fuel}} < \nu$$

Eta or (thermal) fission factor: # of n emitted per thermal neutron absorbed in fuel. Typical value ~ 2.02

$$f = \frac{A_{th}^{fuel}}{A_{th}} = \frac{A_{th}^{fuel}}{A_{th}^{fuel} + A_{th}^{nonFuel}}$$

(Thermal) Fuel utilization factor: fraction of thermal neutrons absorbed *in fuel*. Typical value $\sim .87$

$$p = \frac{A_{th}}{A} = \frac{A - A_{reson}}{A} = 1 - \frac{A_{reson}}{A}$$

Resonance escape probability = fraction of n absorbed *in thermal region*. Typical value ~ 0.8

4 factor formula – Thermal Multiplication Factor

$$\eta_{th} = \frac{P_{th}}{A_{th}^{fuel}} < v$$



Eta or (thermal) fission factor: # of n emitted per thermal neutron absorbed in fuel. Typical value ~ 2.02

- Assume as fuel (F) material UO_2 , with U nuclide density N^F and enrichment \tilde{e}

$$\eta_{th} = \frac{P_{th}}{A_{th}^{fuel}} = \frac{V^F (v \Sigma_{f,th}^F \Phi)}{V^F (\Sigma_{a,th}^F \Phi)} \rightarrow \begin{cases} \Sigma_{f,th}^F = v \tilde{e} N^F \sigma_{f,th}^{235} \\ \Sigma_{a,th}^F = \tilde{e} N^F \sigma_{a,th}^{235} + (1 - \tilde{e}) N^F \sigma_{a,th}^{238} + N^O \sigma_{a,th}^O \\ = \tilde{e} N^F \sigma_{a,th}^{235} + (1 - \tilde{e}) N^F \sigma_{a,th}^{238} + 2 N^F \sigma_{a,th}^O \end{cases}$$

$$\eta_{th} = \frac{v \Sigma_{f,th}^F}{\Sigma_{a,th}^F} = \frac{v \tilde{e} \sigma_{f,th}^{235}}{\tilde{e} \sigma_{a,th}^{235} + (1 - \tilde{e}) \sigma_{a,th}^{238} + 2 \sigma_{a,th}^O} \rightarrow \text{Depends only on fuel properties!}$$

4 factor formula – Fuel Utilization Factor

$$f = \frac{A_{th}^{fuel}}{A_{th}} = \frac{A_{th}^{fuel}}{A_{th}^{fuel} + A_{th}^{nonFuel}}$$



(Thermal) Fuel utilization factor: fraction of thermal neutrons absorbed *in fuel*. Typical value $\sim .87$

- Assume mixture of fuel (F) material UO_2 and moderator (M)
- Homogenous assumption i.e. they “see” the same flux

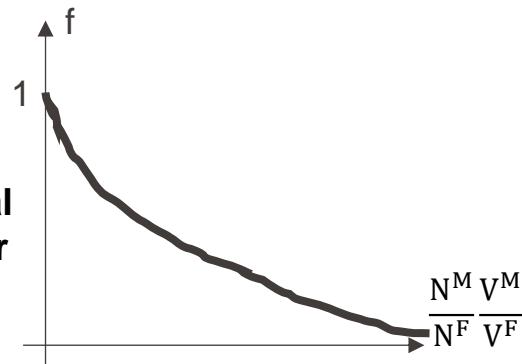
$$f = \frac{A_{th}^{fuel}}{A_{th}} = \frac{V^F \Sigma_{a,th}^F \phi}{V^F \Sigma_{a,th}^F \phi + V^M \Sigma_{a,th}^M \phi}$$

$$\left\{ \begin{array}{l} \Sigma_{a,th}^F = \tilde{\epsilon} N^F \sigma_{a,th}^{235} + (1 - \tilde{\epsilon}) N^F \sigma_{a,th}^{238} + 2 N^F \sigma_{a,th}^0 = N^F \overline{\sigma_{a,th}^F} \\ \text{(where } \overline{\sigma_{a,th}^F} \text{ is an effective average micro-XS for fuel)} \\ \Sigma_{a,th}^{F+M} = N^F \overline{\sigma_{a,th}^F} + N^M \sigma_{a,th}^M \end{array} \right.$$

$$f = \frac{V^F \Sigma_{a,th}^F \phi}{V^F \Sigma_{a,th}^F \phi + V^M \Sigma_{a,th}^M \phi} = \frac{1}{1 + \frac{\sigma_{a,th}^M}{\sigma_{a,th}^F} \cdot \frac{N^M}{N^F} \cdot \frac{V^M}{V^F}}$$



Depends on both material properties and moderator to fuel ratio $\frac{N^M}{N^F} \frac{V^M}{V^F}$!



4 factor formula – Fast Fission Factor

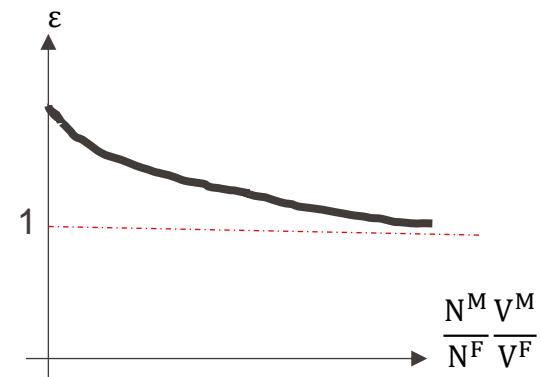
$$\varepsilon = \frac{P}{P_{th}} = \frac{P_{th} + P_{fast}}{P_{th}} = 1 + \frac{P_{fast}}{P_{th}}$$



Fast fission factor: bonus from fast fissions in non-fissile material. Typical value ~ 1.04

$$\varepsilon = 1 + \frac{P_{fast}}{P_{th}} = 1 + \frac{\nu_{fast}^{235} \tilde{\epsilon} N^F \sigma_{f,fast}^{235} + \nu_{fast}^{238} (1 - \tilde{\epsilon}) N^F \sigma_{f,fast}^{238} \phi_{fast}}{\nu_{th}^{235} \tilde{\epsilon} N^F \sigma_{f,th}^{235} \phi_{th}}$$

The ratio on the right depends significantly on the moderator to fuel ratio $\frac{N^M V^M}{N^F V^F}$ between 0.02 and 0.3!

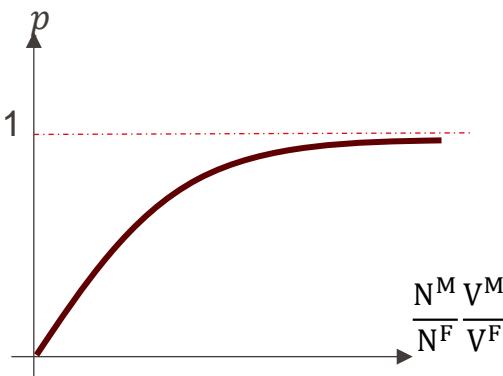


$$p = \frac{A_{th}}{A} = \frac{A - A_{reson}}{A} = 1 - \frac{A_{reson}}{A}$$



Resonance escape probability = fraction of n absorbed *in thermal region*. Typical value ~ 0.8

$$\begin{aligned} p &\approx \exp\left(-\frac{V^F N^F (1 - \tilde{e})}{V^M \xi^M \Sigma_s^M} I\right) \\ &= \exp\left(-\frac{V^F N^F (1 - \tilde{e})}{N^M V^M \xi^M \sigma_s^M} I\right) \end{aligned}$$

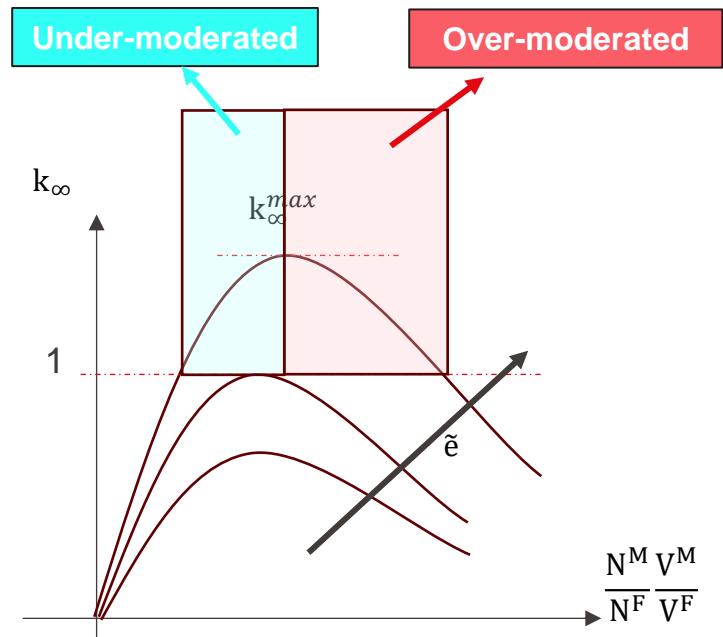


- I is the resonance integral; it describes the effectiveness of the moderator in slowing down the neutrons past the resonance region.
- Σ_s^M is the scattering cross section of the moderator.
- $N^F(1 - \tilde{e})$ is the number density of the fertile material (U238).
- ξ^M is the average logarithmic energy decrease per collision (see later slides)
- p is also a function of fuel to moderator ratio \rightarrow tends toward 1 as moderator ratio increases.

- In our simplified model of a thermal reactor, we found that k_{∞} is a function of material properties (XS, densities, enrichment) and of moderator to fuel ratio!

$$k_{\infty} = \epsilon \left(\frac{N^M V^M}{N^F V^F} \right) \times \eta_{\text{th}} \times f \left(\frac{N^M V^M}{N^F V^F} \right) \times p \left(\frac{N^M V^M}{N^F V^F} \right)$$

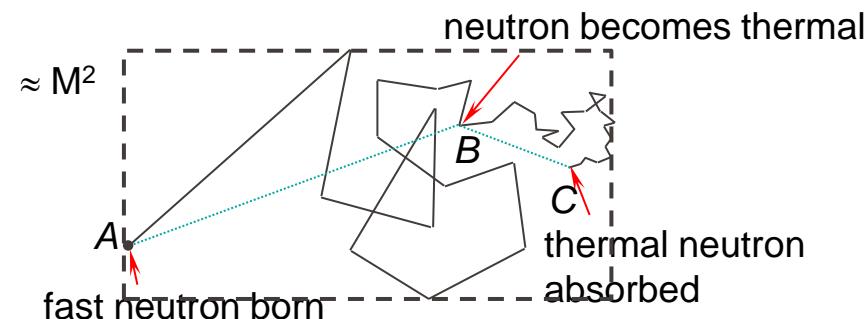
- As a function of moderator to fuel ratio, k_{∞} increases, reaches a maximum and then decreases (different effects at play).
- To have a critical reactor $k_{\infty} > 1$, so there is a lower limit enrichment we can allow in a given reactor (that for which $\max k_{\infty} = 1$).
- LWR are operated typically in **under-moderated regions**: when moderator density decreases, the multiplication factor decreases.



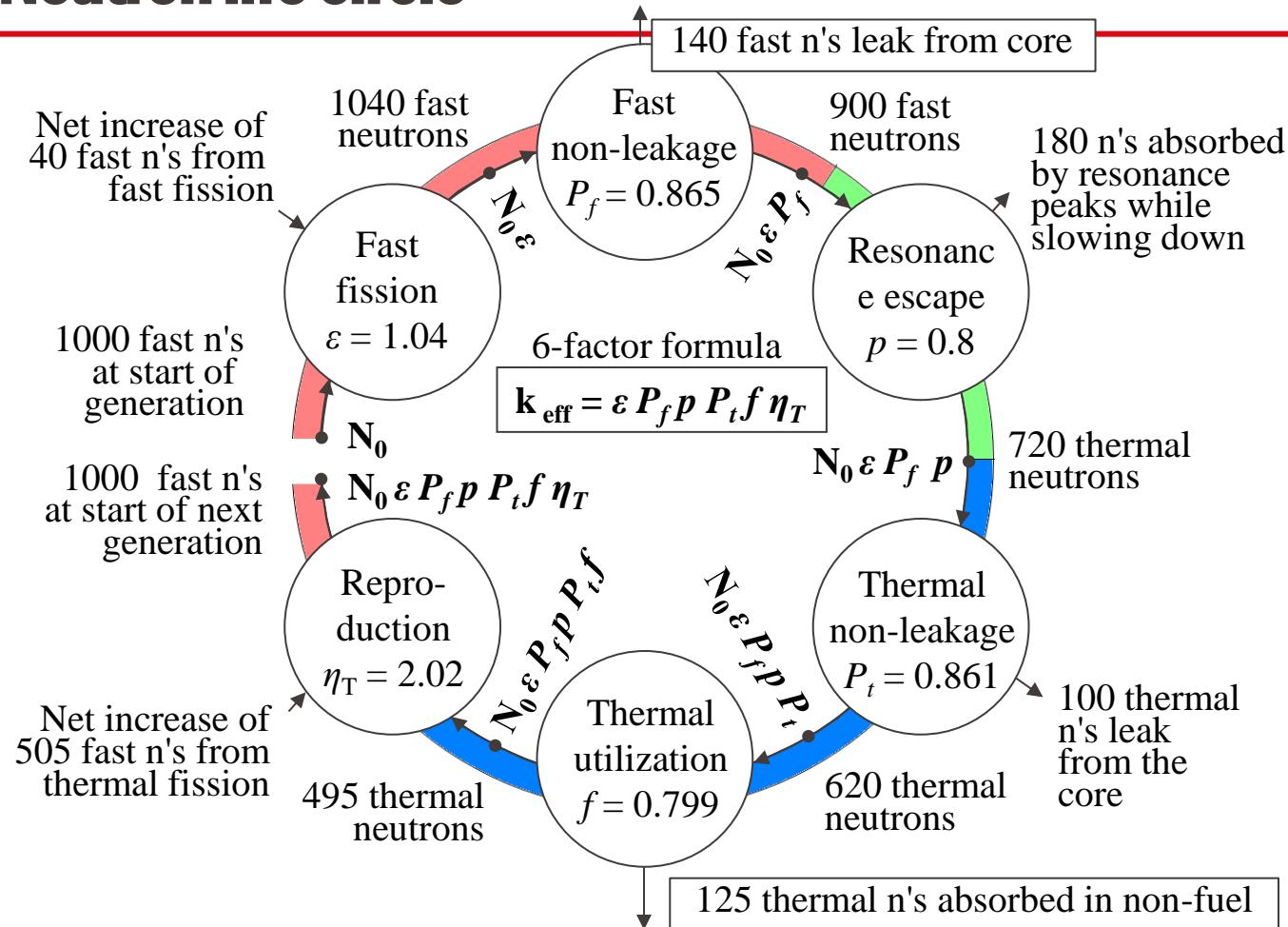
- For a reactor of finite size, one needs to consider the fraction of neutrons that leaks out:
 - Non-leakage probabilities: P_f (for fast neutrons), P_{th} (for thermal)
- Thus, the effective multiplication factor is: $k_{eff} = k_{\infty} P_f P_{th} = \varepsilon \eta_{th} f p P_f P_{th}$
- With respect to one-group neutron balance one finds

$$P_{NF} = P_f P_{th} \approx \frac{1}{1 + M^2 B^2}$$

$$k_{eff} \approx \frac{k_{\infty}}{1 + M^2 B^2}$$



Neutron life circle

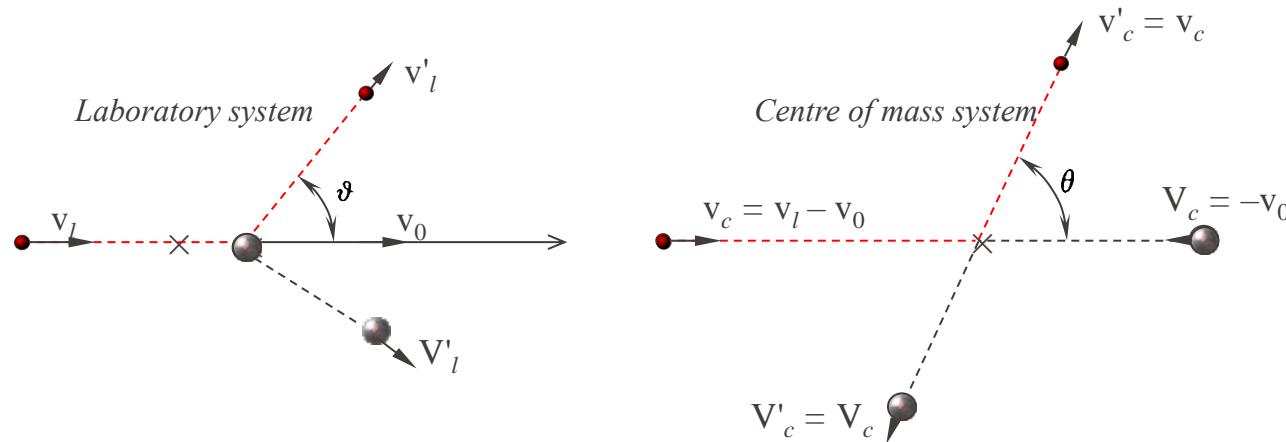


- A thermal reactor has neutrons between ~ 2 MeV and ~ 0.01 eV
 - One needs to study how neutrons change E on average across this range.
 - Slowing down process determines the “thermal -neutron source”

→ One needs to determine the neutron energy spectrum (*flux as a function of space and energy*) for evaluating different reaction rates.

→ ***In LWRs, one needs to choose a material to slow down the neutron efficiently, i.e. with as little absorption as possible.***

- Most important slowing-down mechanism: elastic scattering by moderator nuclei
 - Inelastic scattering also plays a role, but only for fast neutrons ($E \geq 1$ MeV)



- Typically, one considers the c-mass system where the velocity before and after collision remain the same, only direction changes (derivation not discussed).

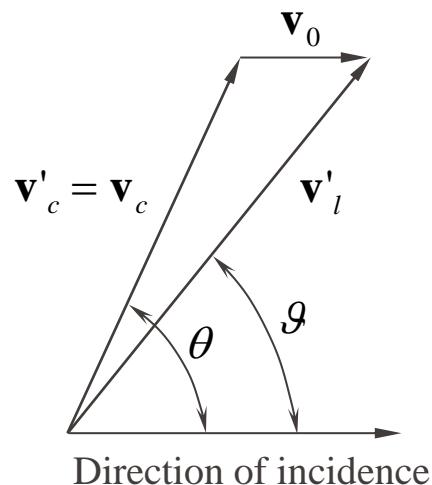
Using the c-mass system one obtains (skipping derivation) the fractional energy loss as:

$$\frac{\Delta E_l}{E_l} = 1 - \frac{E'_l}{E_l} = 1 - \frac{(v'_l)^2}{(v_l)^2} = 1 - \frac{A^2 + 2A \cos \theta + 1}{(A + 1)^2} = (1 - \alpha) \frac{1 - \cos \theta}{2}$$

where the new parameter $\alpha = \left(\frac{A-1}{A+1}\right)^2$ depends only on material mass number.

NOTE that:

- When $\theta = 0$, there is the minimum energy loss i.e. $\frac{\Delta E_l}{E_l} = 0$ or $E'_l = E_l$
- For $\theta = \pi$, there is the maximum energy loss i.e. $\frac{\Delta E_l}{E_l} = (1 - \alpha)$ or $E'_l = \alpha E_l$

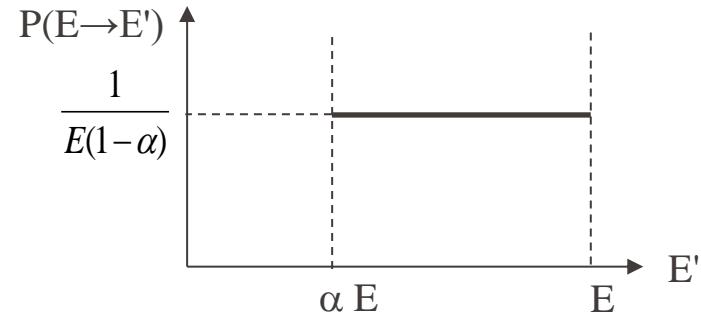


- Probability distribution function $P(E \rightarrow E')$:

$P(E \rightarrow E')dE'$ is the probability that a neutron with the laboratory energy E will have after collision an energy between E' and $E' + dE'$

- In case of isotropic scattering in the center-of-mass system:

$$P(E \rightarrow E') = \begin{cases} \frac{1}{E(1-\alpha)} & \alpha E < E' < E \\ 0 & 0 < E' < \alpha E \end{cases}$$



- For hydrogen: $P(E \rightarrow E') = \frac{1}{E}$

- Average energy loss depends on material (α) and energy E

$$\langle \Delta E \rangle = \int_0^{\infty} (E - E') \cdot P(E \rightarrow E') dE' = \int_{\alpha E}^E dE' \frac{E - E'}{(1 - \alpha)E} = \frac{1 - \alpha}{2} E$$

- One finds that **the average logarithmic energy loss per collisions (slow down decrement or ξ)** depends not on energy but only on A !

$$\xi = \int_{\alpha E}^E \ln\left(\frac{E}{E'}\right) P(E \rightarrow E') dE' \quad \Rightarrow \quad \xi = \frac{1}{E(1-\alpha)} \int_{\alpha E}^E \ln\left(\frac{E}{E'}\right) dE'$$

$$x = E'/E \quad \Rightarrow \quad \xi = \frac{1}{1 - \alpha} \int_1^{\alpha} \ln x \, dx = 1 + \frac{\alpha}{1 - \alpha} \ln \alpha$$

For $A > 10$

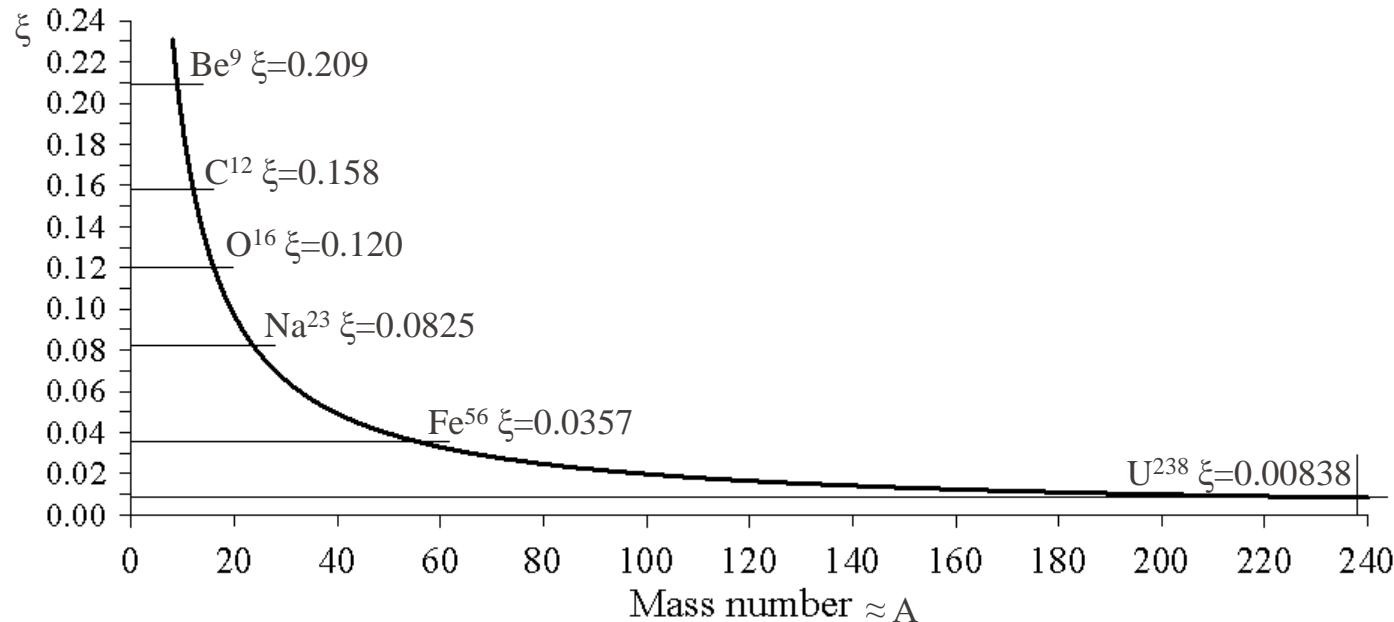
$$\xi \approx \frac{2}{A + \frac{2}{3}}$$

- For a mixture of nuclides:

$$\bar{\xi} = \frac{1}{\Sigma_s} \sum_i \Sigma_{si} \int_0^{\infty} \ln\left(\frac{E}{E'}\right) \cdot P_i(E \rightarrow E') dE' = \frac{1}{\Sigma_s} \sum_i \Sigma_{si} \xi_i$$

- So for instance for water:

$$\overline{\xi_{H_2O}} = \frac{\Sigma_{sH} \xi_H + \Sigma_{so} \xi_O}{\Sigma_{sH} + \Sigma_{so}} = \frac{2\sigma_{sH} \xi_H + \sigma_{so} \xi_O}{2\sigma_{sH} + \sigma_{so}}$$



Not just the slow down decrement qualifies/quantifies a moderator:

- **Slow down decrement ξ** determines the minimum number of collisions to thermalize $N = \ln(2 \text{ MeV} / 0.025 \text{ eV}) / \xi$
- A good moderator must also have Σ_s (thus its density is important). This means that the **macroscopic slowing-down power** = $\xi \Sigma_s$ **must be large!**
- A good moderator must also not absorb much. **Moderating Ratio** = $\xi \Sigma_s / \Sigma_a$ **must be large!**

TABLE 2
Moderating Properties of Materials

Material	ξ	Number of Collisions to Thermalize	Macroscopic Slowing Down Power	Moderating Ratio
H ₂ O	0.927	19	1.425	62
D ₂ O	0.510	35	0.177	4830
Helium	0.427	42	9×10^{-6}	51
Beryllium	0.207	86	0.154	126
Boron	0.171	105	0.092	0.00086
Carbon	0.158	114	0.083	216